

Física I

LEIC-T 2021-2022

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6ª Aula

Momento de Força e variação do Momento Angular;
A equação para o pêndulo simples e para o pêndulo físico;
Equilíbrio estático;
Forças centrais e Conservação do Momento Angular;
Energia potencial efetiva e órbitas de transferência;
Leis de Kepler do Movimento Planetário.

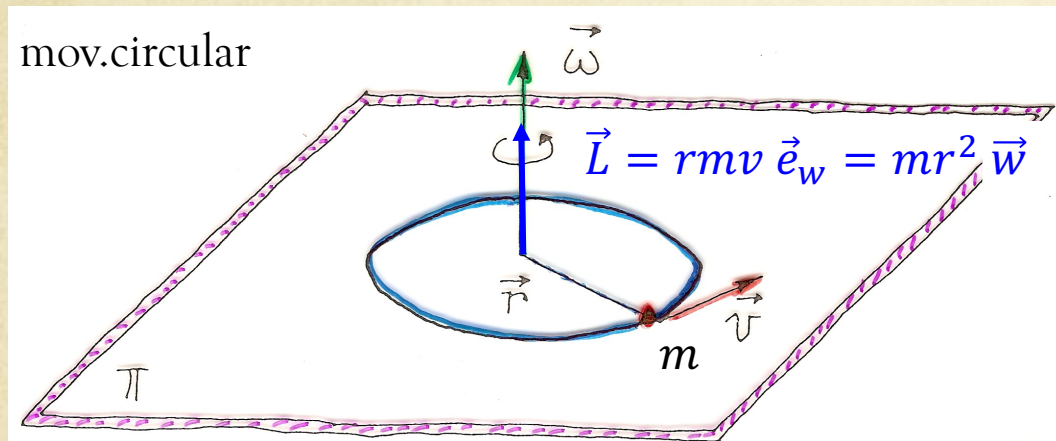
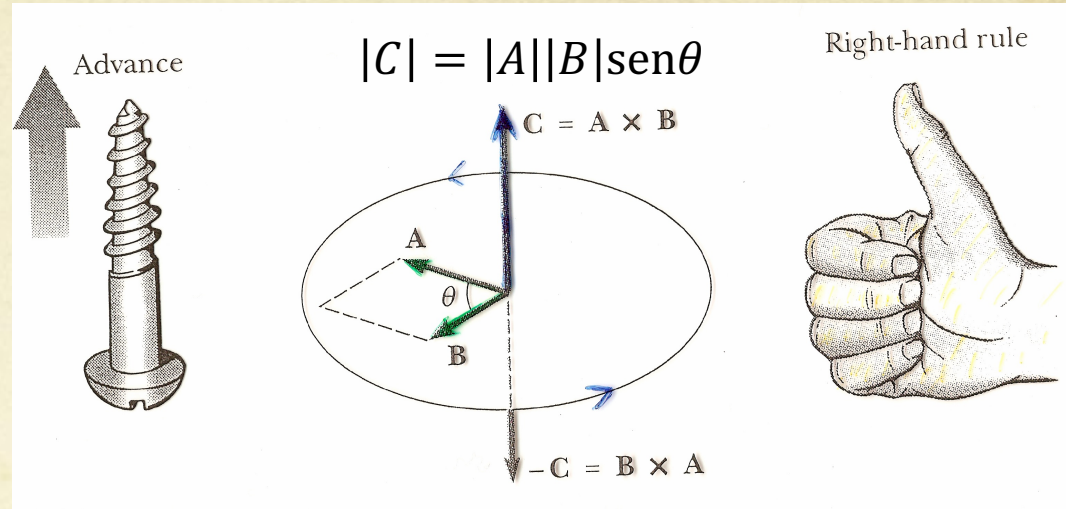
*“A viagem real de descoberta não consiste em procurar novas paisagens mas
Marcel Proust (1871-1922) em ter uma nova visão!”*

O Momento Angular (recapitulação)

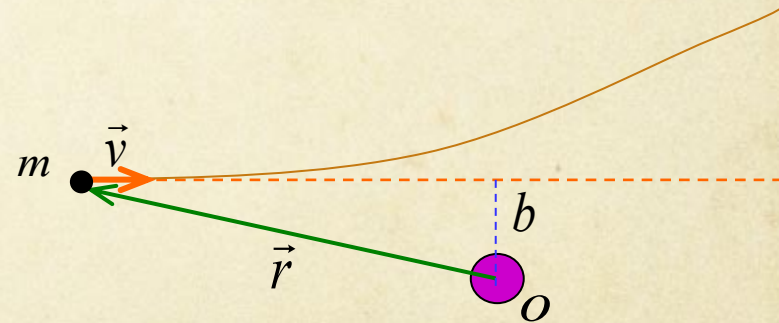
$$\vec{L} = \vec{r} \times \vec{p} \quad [\text{kgm}^2/\text{s}]$$

$$|\vec{L}| = |\vec{r}_i \times \vec{p}_i| \equiv r_i p_i \text{sen}\theta$$

$$\vec{L} \perp \vec{r}, \vec{p}$$



$$\vec{L} = \vec{r} \times m \vec{v} = m v b \vec{e}_\perp$$



Rotação no plano (corpo pontual)

n massas pontuais:

$$\vec{L}_O = \sum \vec{r}_{iO} \times \vec{p}_i$$

Variação do Momento Angular

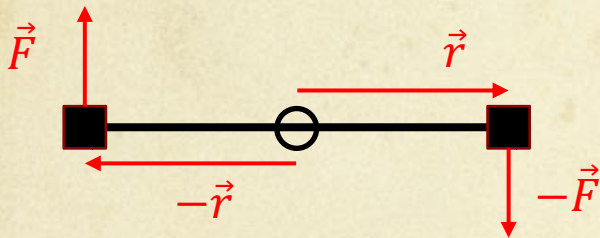
$$\vec{L} = \sum \vec{r}_i \times \vec{p}_i$$

$$\frac{d\vec{L}}{dt} = \sum \vec{r}_i \times \vec{F}_i = \vec{N} = \vec{\tau}$$

(Momento de Força, *torque*)
[N.m]

Binário:

(centro fixo)



$$\sum \vec{F}_i = 0 \quad \text{mas}$$

$$\sum \vec{r}_i \times \vec{F}_i = rF\vec{e}_{\otimes} + rF\vec{e}_{\otimes} = 2rF\vec{e}_{\otimes}$$

$$\vec{N} = \vec{\tau} = 2rF\vec{e}_{\otimes} \neq 0 \Rightarrow d\vec{L} = 2rFdt\vec{e}_{\otimes} : \text{ Existe Rotação em torno do eixo (centro)}$$

Momento do Binário (ou simplesmente, Binário!)



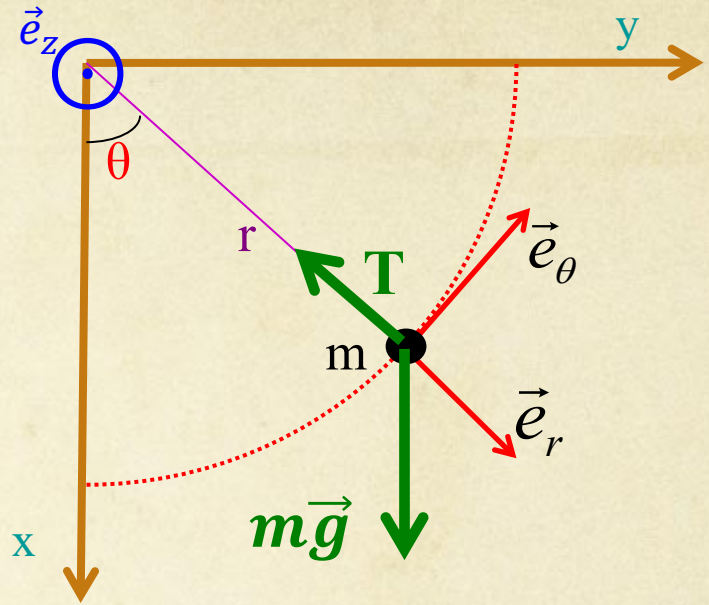
Pêndulo simples revisitado

$$\vec{N} \equiv \frac{d\vec{L}}{dt}$$

$$\vec{L} = \vec{r} \times \vec{p} = mr^2 \dot{\theta} \vec{e}_z$$

$$\vec{N} = \vec{r} \times (\vec{T} + m\vec{g}) = mr^2 \ddot{\theta} \vec{e}_z \Leftrightarrow$$

$$-r mg \sin \theta \vec{e}_z = mr^2 \ddot{\theta} \vec{e}_z \Leftrightarrow \ddot{\theta} + \frac{g}{r} \sin \theta = 0$$



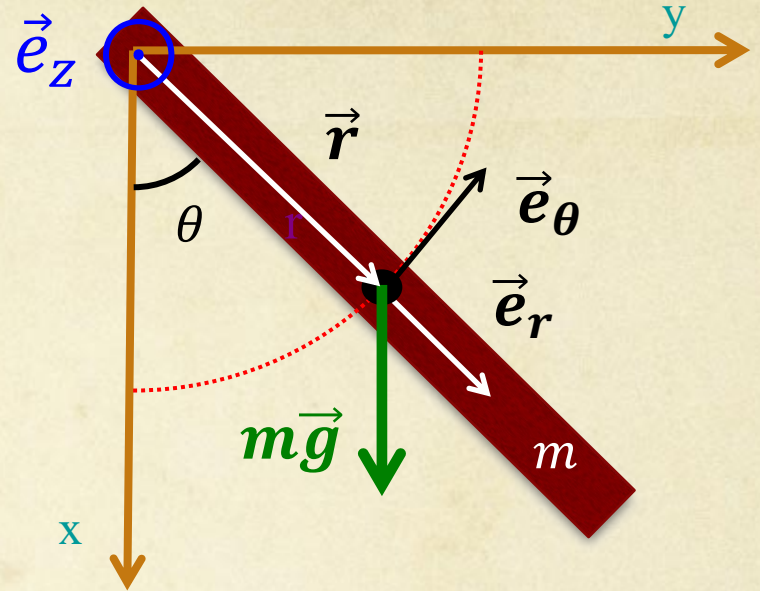
Pêndulo Físico (barra presa num extremo)

$$\vec{N} \equiv \frac{d\vec{L}}{dt}$$

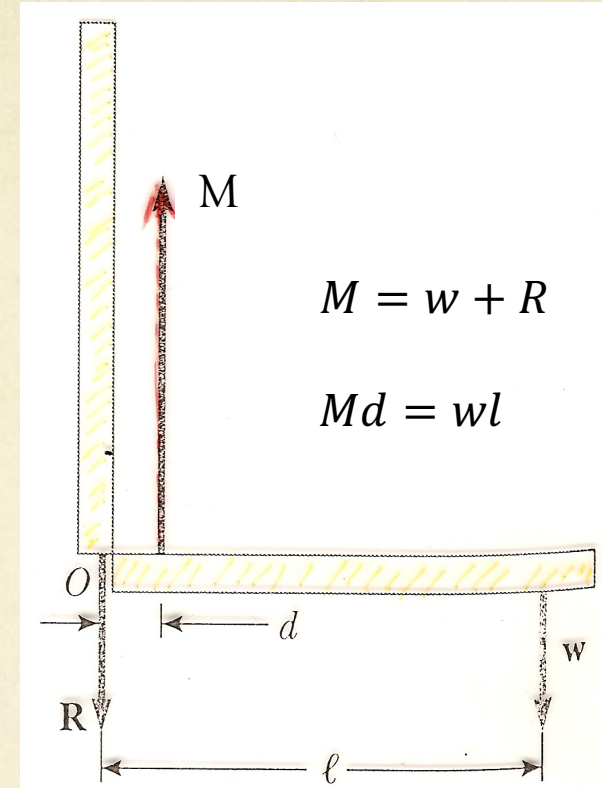
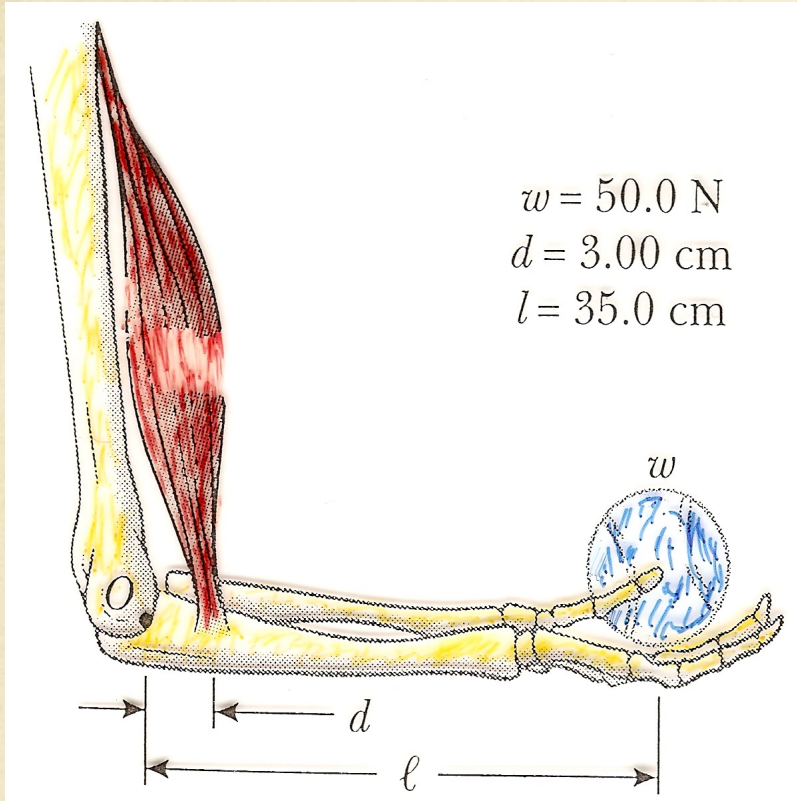
$$\vec{L} = I_O \vec{\omega} = \left(\frac{ml^2}{12} + m \left(\frac{l}{2} \right)^2 \right) \dot{\theta} \vec{e}_z$$

$$\vec{N} = \vec{r} \times m \vec{g} = \frac{ml^2}{3} \ddot{\theta} \vec{e}_z \Leftrightarrow$$

$$-\frac{l}{2} mg \text{ sen } \theta \vec{e}_z = \frac{ml^2}{3} \ddot{\theta} \vec{e}_z \Leftrightarrow \ddot{\theta} + \frac{3g}{2l} \text{ sen } \theta = 0$$



Equilíbrio estático



$$\sum \vec{F} = 0 \quad \text{e} \quad \sum \vec{N} = 0$$

Forças centrais

$$\vec{F} \parallel \vec{r}$$

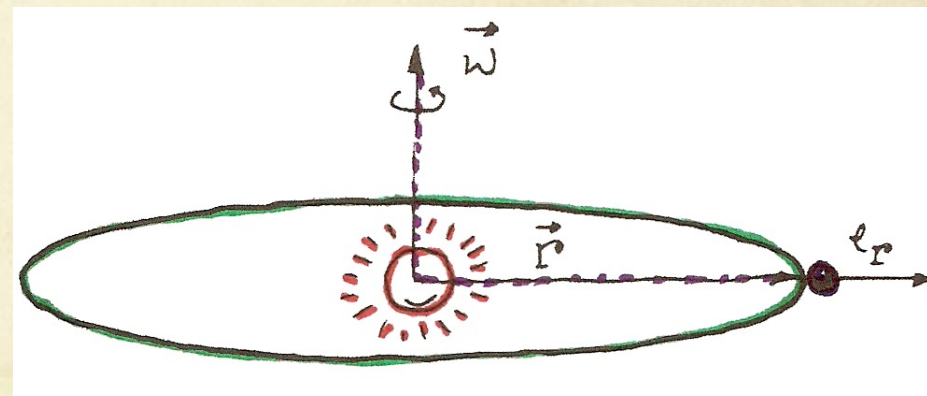
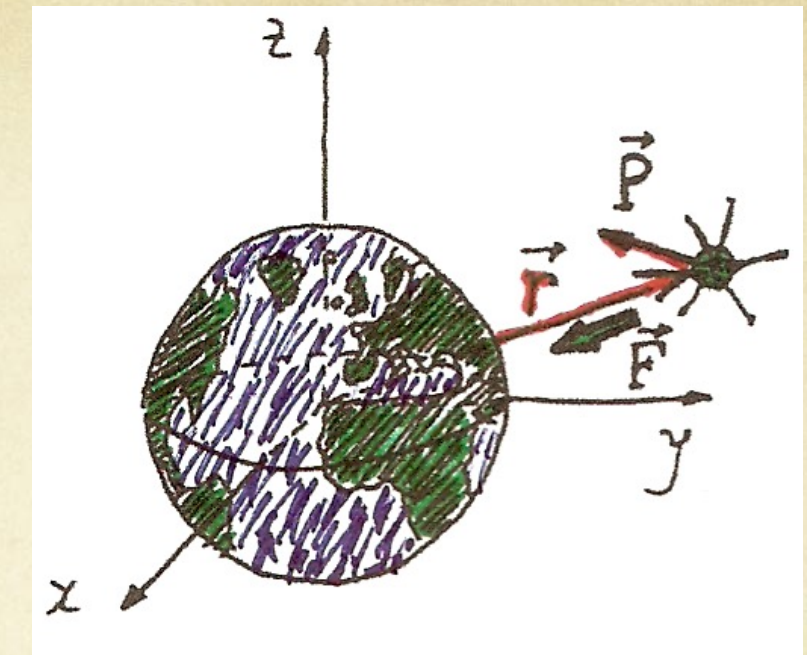
$$\frac{d\vec{L}}{dt} = \vec{r} \times \vec{F} = 0$$

$$\vec{L} = \text{cte} = mr^2\dot{\theta}\vec{e}_w$$

$$E_C = \frac{1}{2}m\dot{r}^2 + \frac{1}{2}mr^2\dot{\theta}^2 \quad (*)$$

$$E = E_C + E_P$$

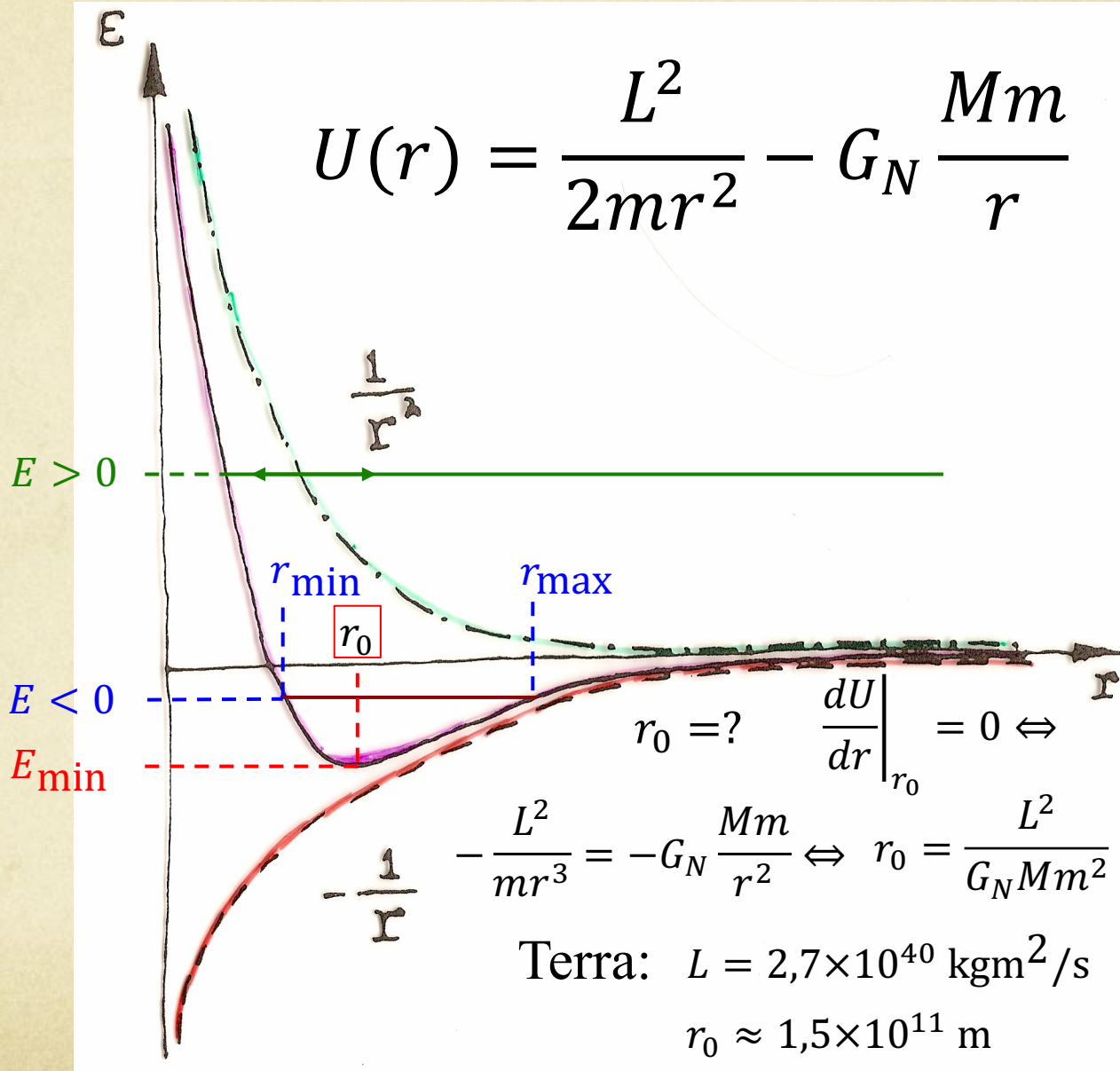
$$E = \frac{1}{2}m\dot{r}^2 + \frac{L^2}{2mr^2} - G_N \frac{Mm}{r}$$



$$E_{P \text{ efectiva}} = U(r)$$

$$(*) \vec{v} = \dot{r}\vec{e}_r + r\dot{\theta}\vec{e}_\theta \Leftrightarrow v^2 = \dot{r}^2 + r^2\dot{\theta}^2$$

Energia potencial efetiva



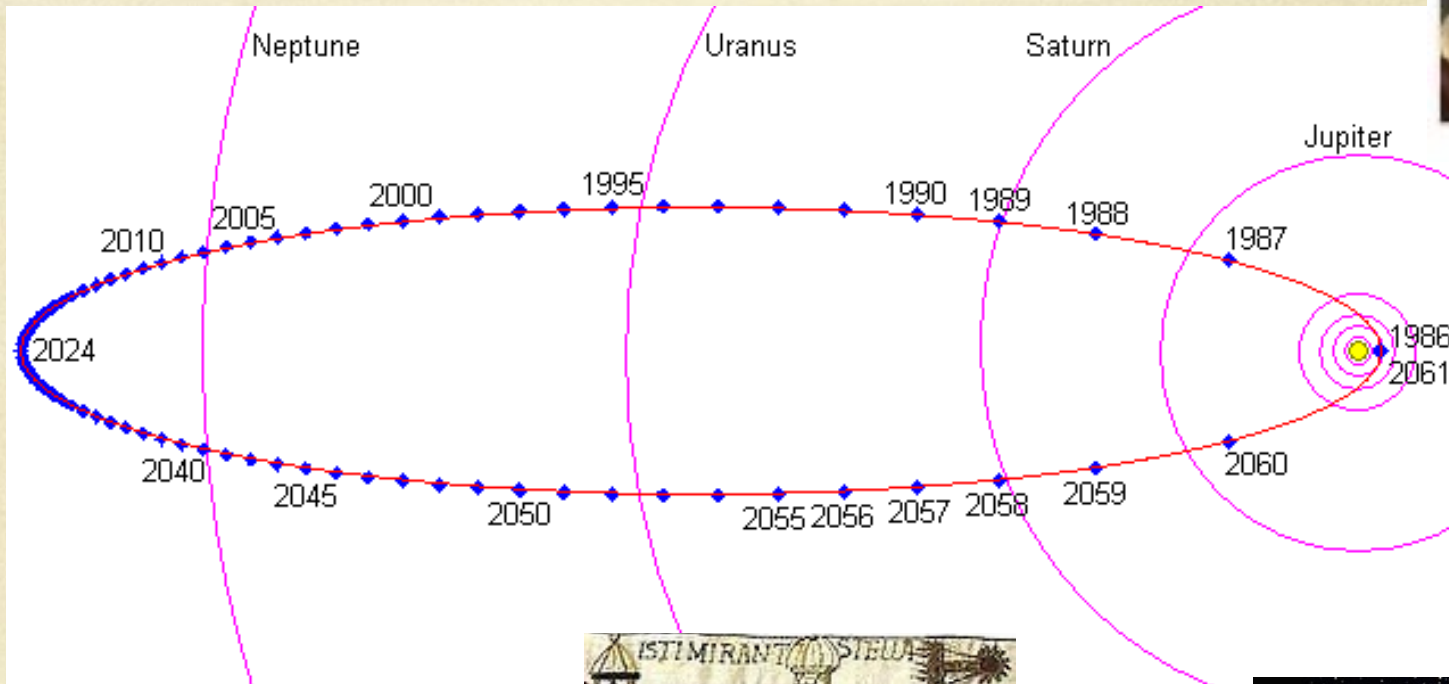
O cometa Halley

T ~ 75 anos

Edmond Halley



1656-1742



164 AC



1066



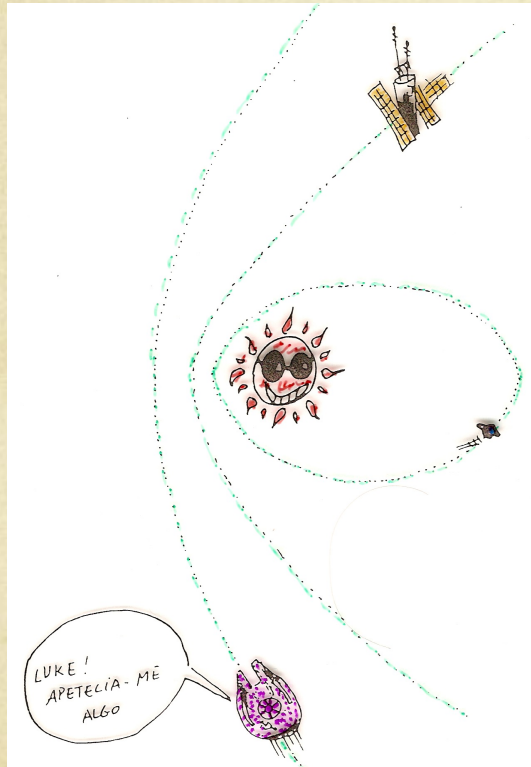
1986

Órbitas de transferência

$$E, \vec{L} = \text{cte}$$

1 → 3

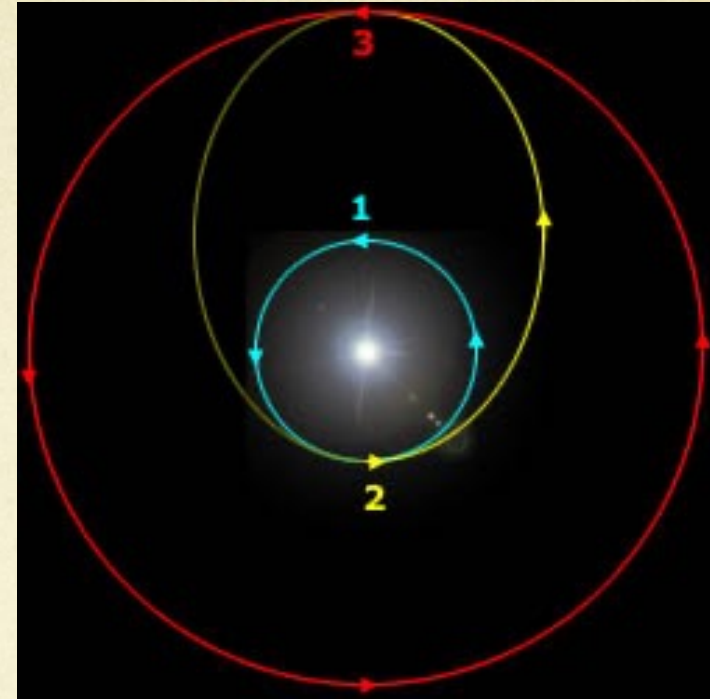
$$\Delta E = \frac{1}{2} G_N M m \left(\frac{1}{r_1} - \frac{1}{r_3} \right) = \frac{G_N M m}{2 r_1 r_3} (r_3 - r_1)$$



1 → 2 + 2 → 3

$$\Delta v_{12} = v_{2i} - v_1 = \frac{L_2 - L_1}{m r_1}$$

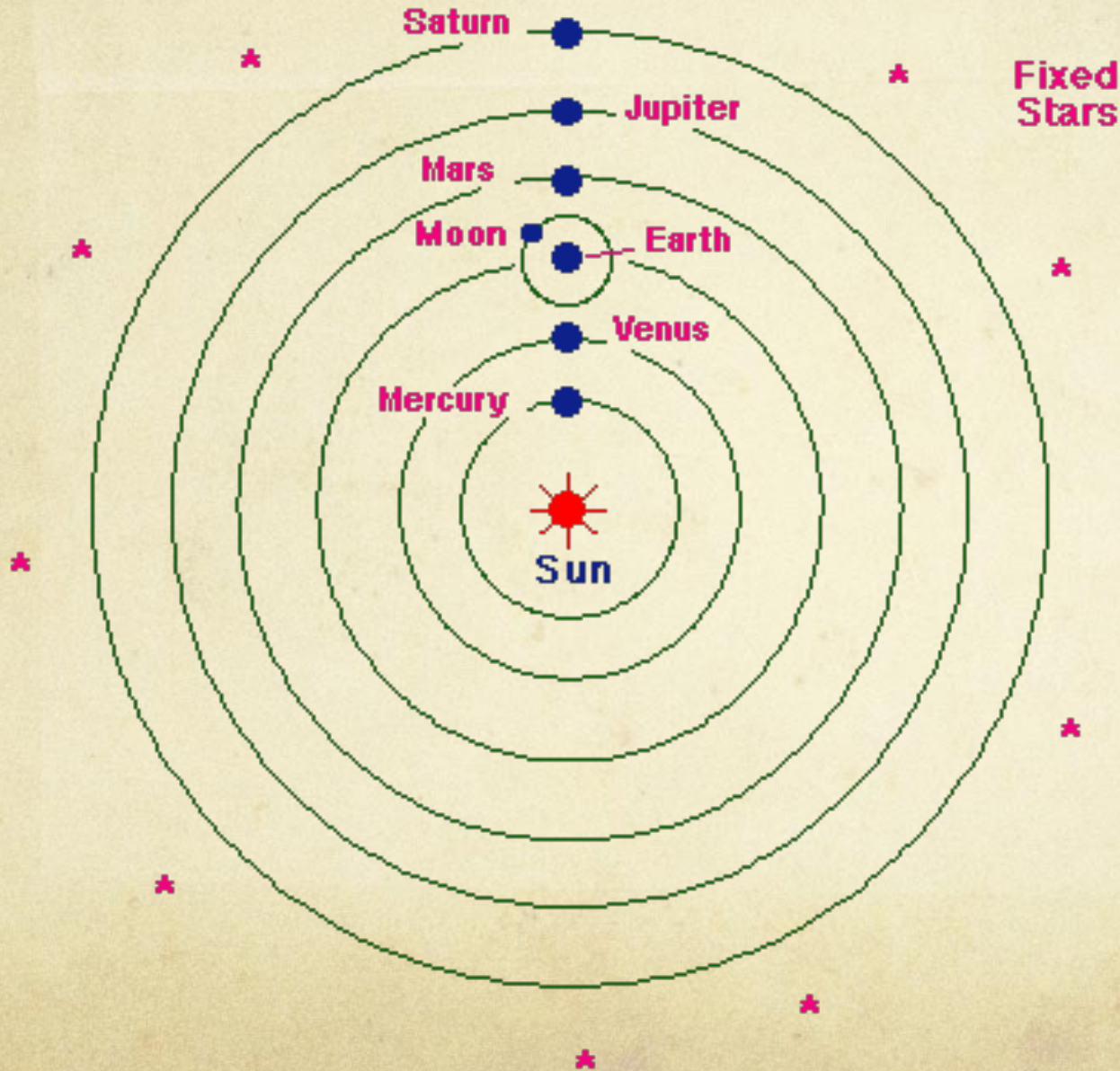
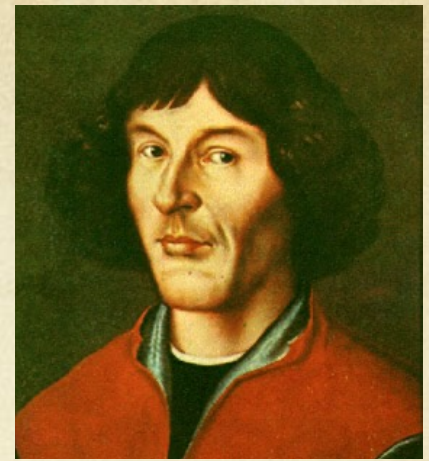
$$\Delta v_{23} = v_3 - v_{2f} = \frac{L_3 - L_2}{m r_3}$$



$$L_2 = L_{\text{transfer}}$$

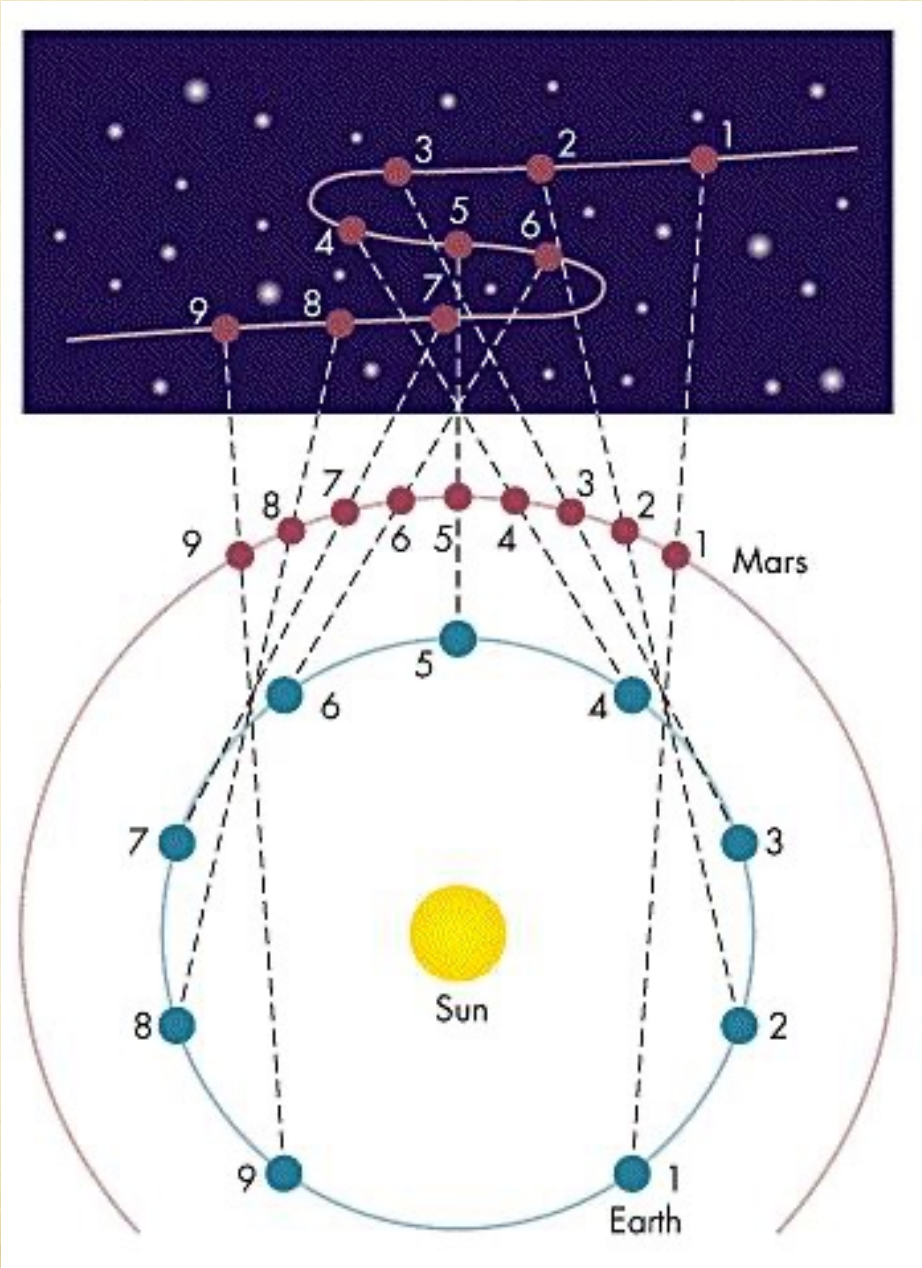
com
$$L_{\text{transfer}}^2 = 2m^2 G_N M \left(\frac{r_1 r_3}{r_1 + r_3} \right)$$

Copérnico (1473-1543)



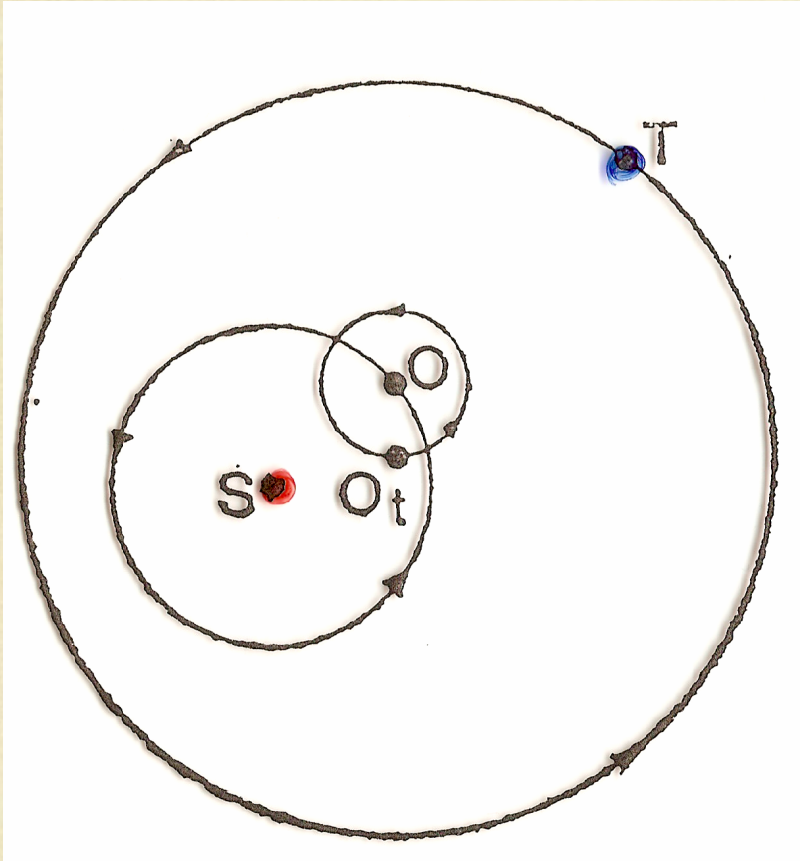
De Revolutionibus
Orbium Caelestium

O Movimento retrógrado

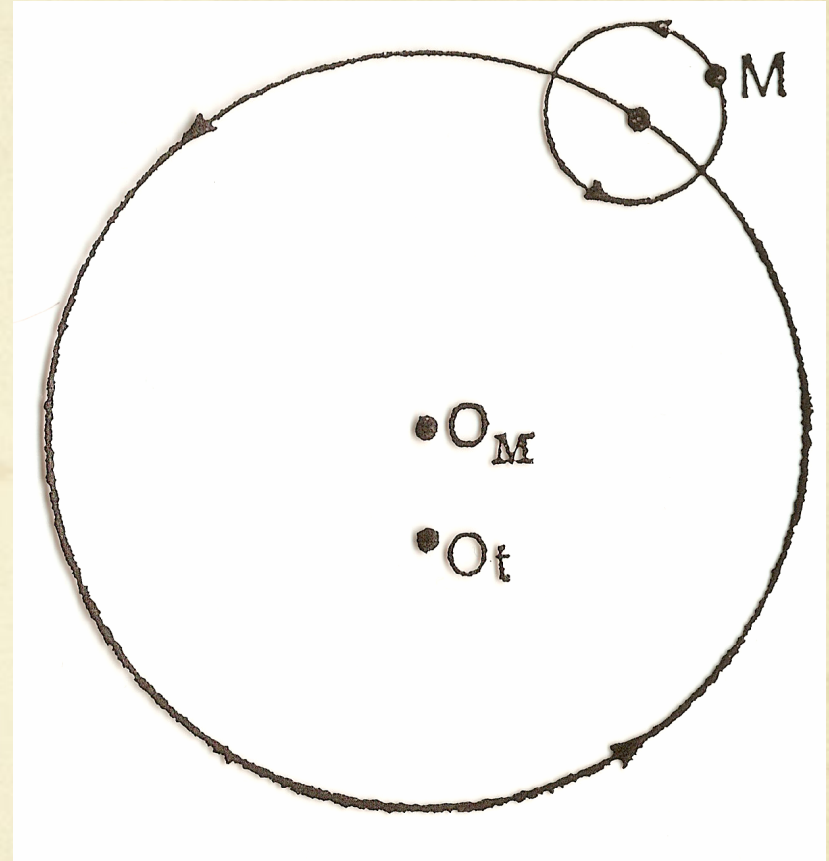


Mas não chega !

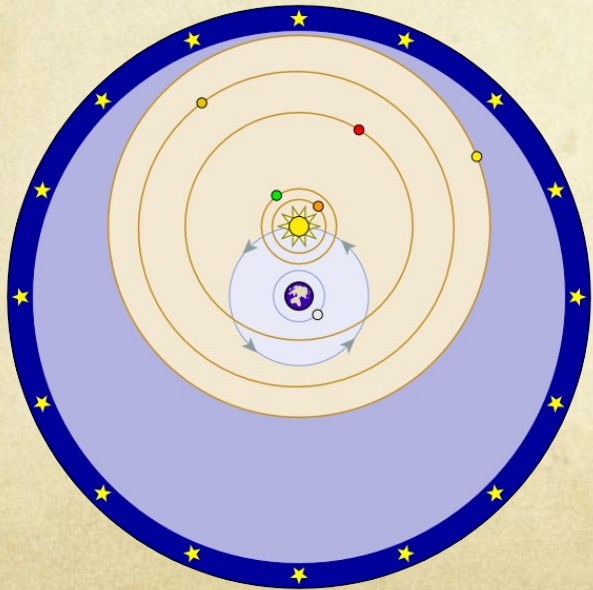
Órbita da Terra



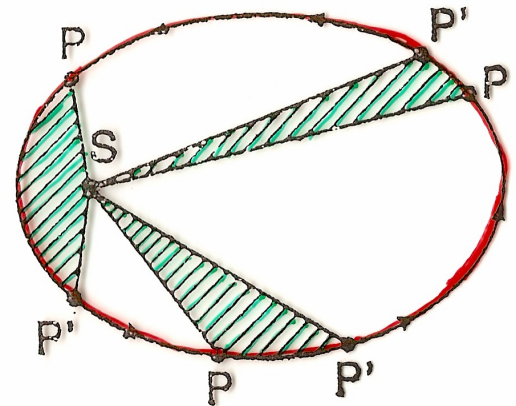
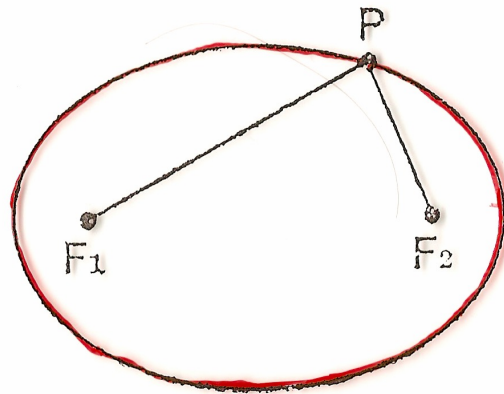
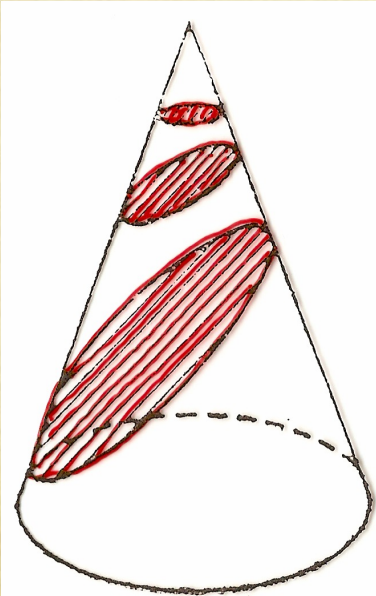
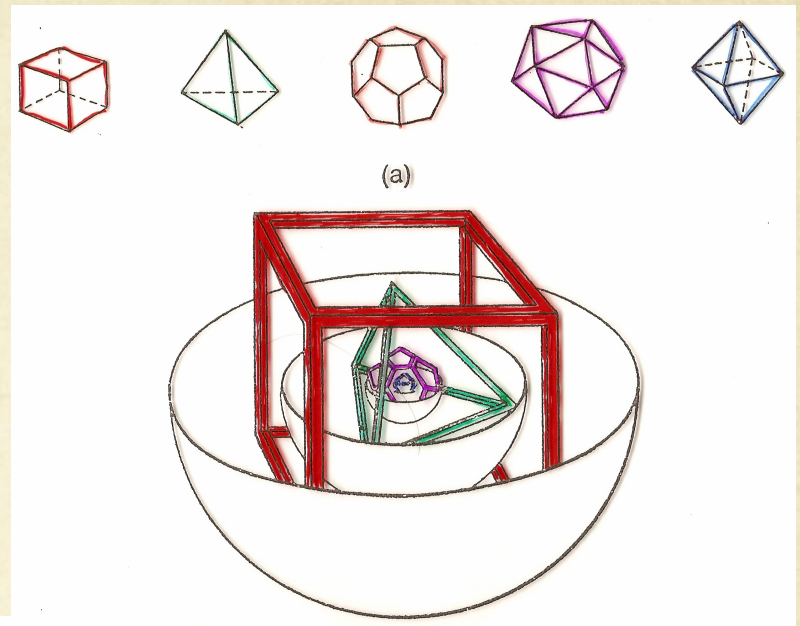
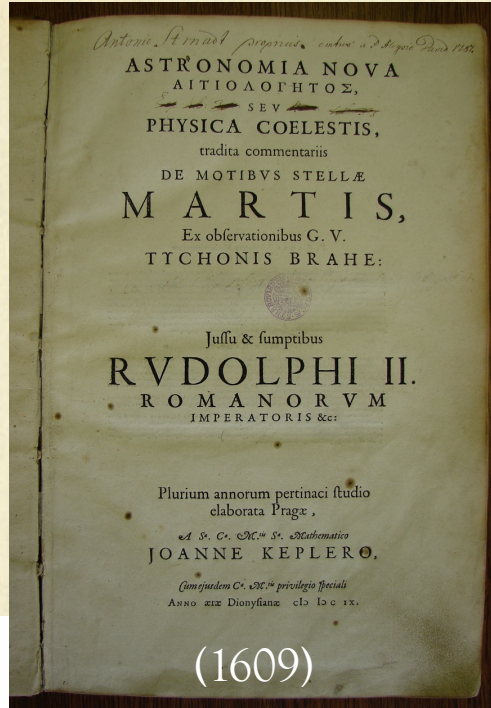
Órbita de Marte



Tycho Brahe (1546-1601)



Johannes Kepler (1571-1630)

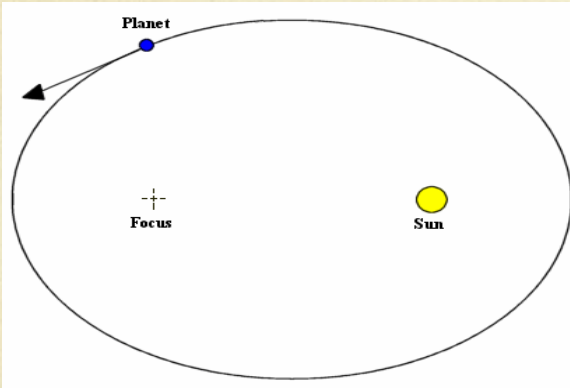


Leis de Kepler

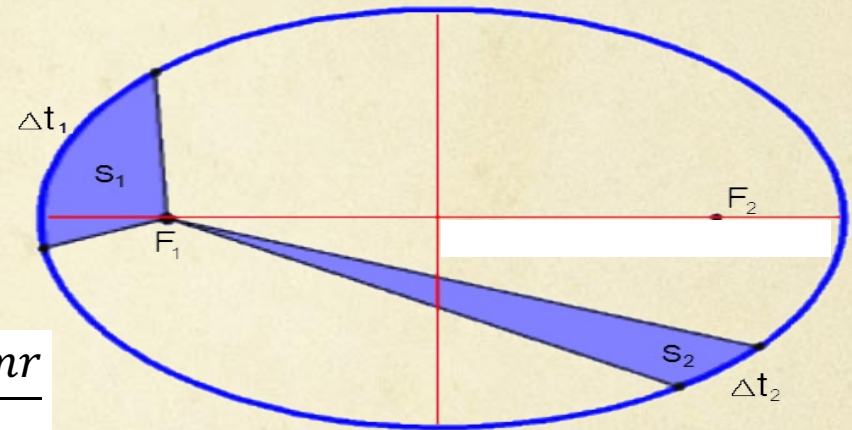
$$\vec{F} \parallel -\vec{r} \quad \vec{r} \times m\vec{v} = \text{cte} \quad \vec{v} \perp \vec{r}$$



1ª Lei – Movimento num plano; Órbitas elípticas (Sol no foco)



2ª Lei – Áreas iguais em tempos iguais



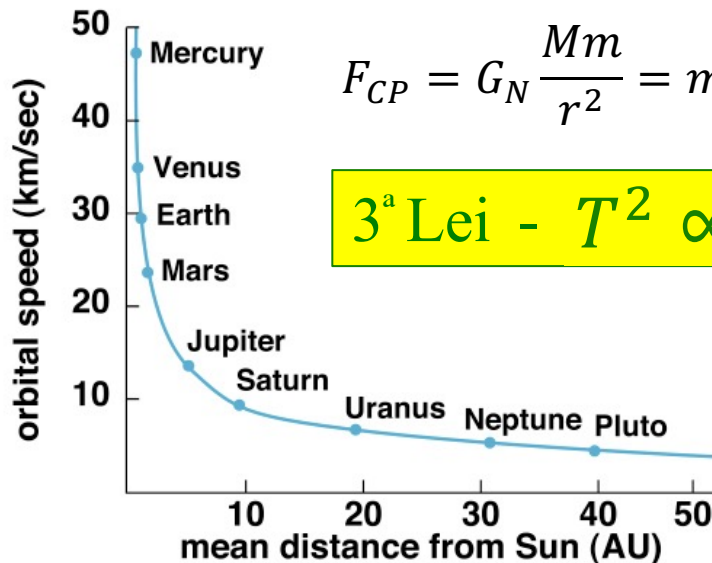
(para órbitas aprox. circulares temos:)

$$v = \omega r \quad L = \text{cte} = mvr = mr^2 \frac{d\theta}{dt}$$

Sist.solar:

$$\frac{L}{m} dt = r \cdot r d\theta = 2 \text{ área}$$

$$T_T = \frac{2\pi}{\sqrt{G_N M_\odot}} r_T^{3/2} = 5,44 \times 10^{-10} r_T^{3/2}$$



3ª Lei - $T^2 \propto r^3$

$$F_{CP} = G_N \frac{Mm}{r^2} = m\omega^2 r = \frac{4\pi^2 m r}{T^2}$$